

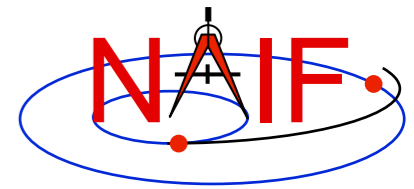
Introduction au logiciel

SPICE

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Qu'est-ce que c'est?



- Logiciel développé et maintenu par le **Navigation and Ancillary Information Facility (NAIF)** situé au Jet Propulsion Laboratory (JPL) à Pasadena (Californie)
- Utilisé pour l'ensemble des projets NASA d'exploration planétaire (Missions martiennes, Cassini, Deep Impact, Messenger, Juno)
- Support à certaines missions «astrophysiques» (Hubble, Spitzer, Kepler, etc.)
- Utilisé pour de nombreuses missions européennes (Smart-1, Mars/Venus Express, Rosetta, Huygens, etc.), russes (Mars 96), japonaises (Hayabusa, SELENE), indienne (Chandrayaan-1)
- Utilisé pour le NASA DSN

A quoi ça sert?

Reference frames

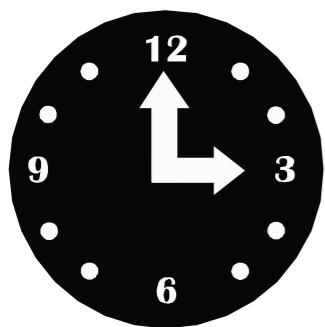
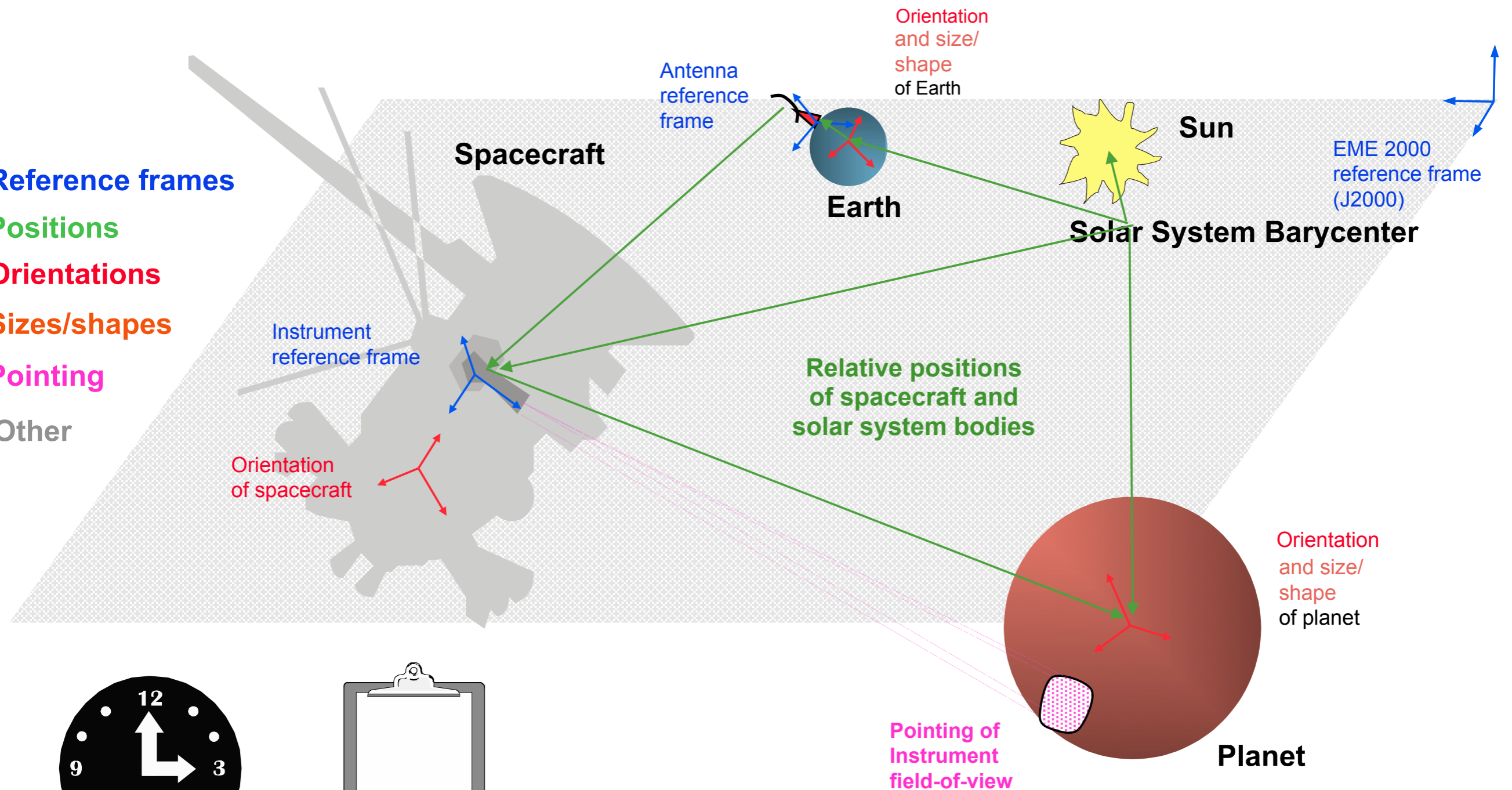
Positions

Orientations

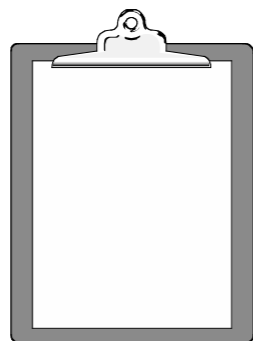
Sizes/shapes

Pointing

Other



Time Conversion
Calculations



Logs of Commands
and Events

Overview of SPICE

Quoi est où par rapport à quoi (à quel moment et dans quel repère)?

Ça veut dire quoi?

S

Spacecraft

P

Planet

I

Instrument

C

C-matrix

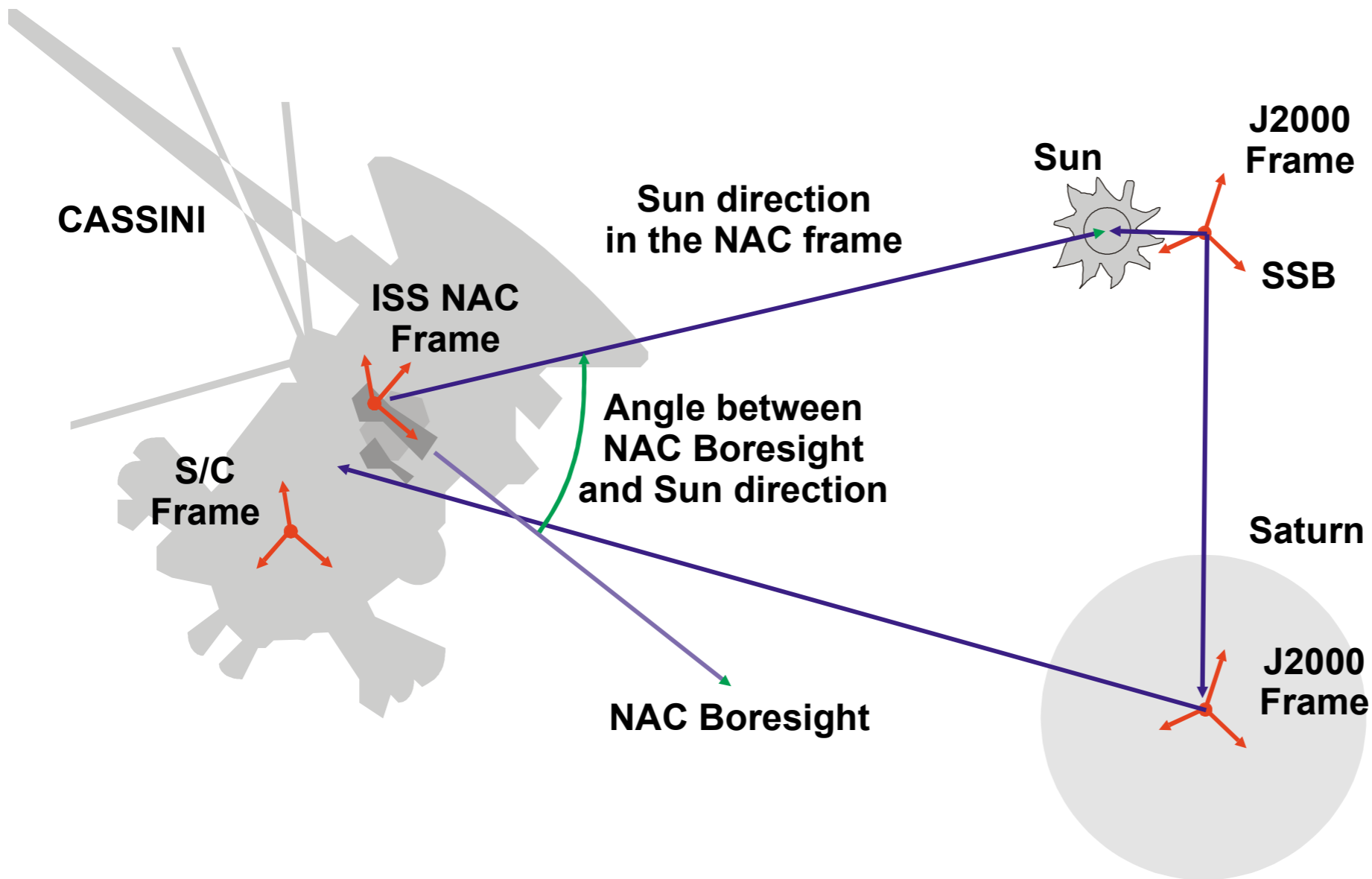
E

Events



* Coined by Dr. Hugh Kieffer, USGS Astrogeology Branch, Flagstaff AZ

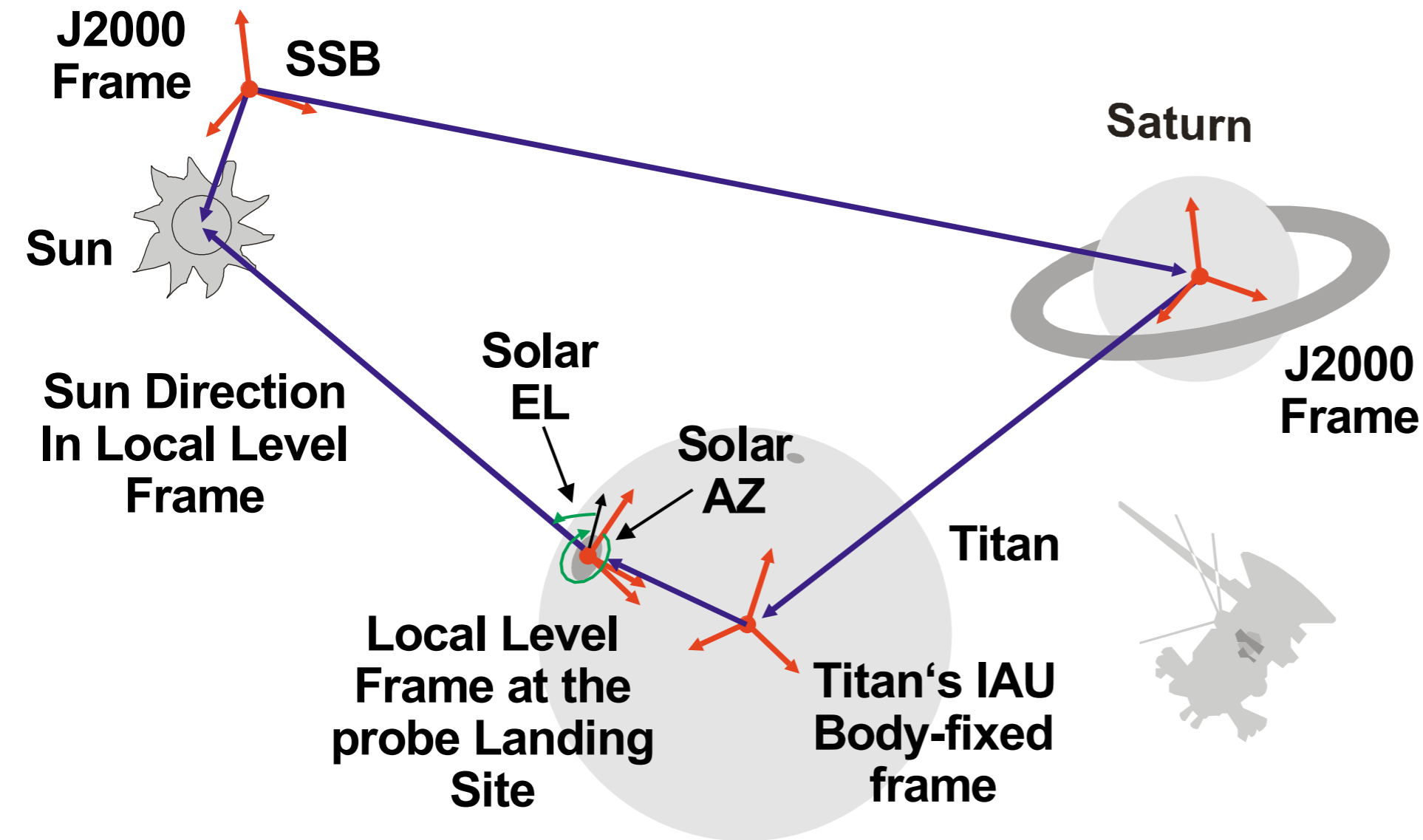
Example 1



Check whether the angle between the camera boresight and the direction to the Sun is within the allowed range:

```
CALL SPKPOS( 'SUN', ET, 'CASSINI_ISS_NAC', 'LT+S', 'CASSINI', SUNVEC, LT )
ANGLE = VSEP( NAC_BORESIGHT_nac, SUNVEC )
IF ( ANGLE .LE. CONSTRAINT ) WRITE(*,*) 'WE ARE IN TROUBLE!'
```

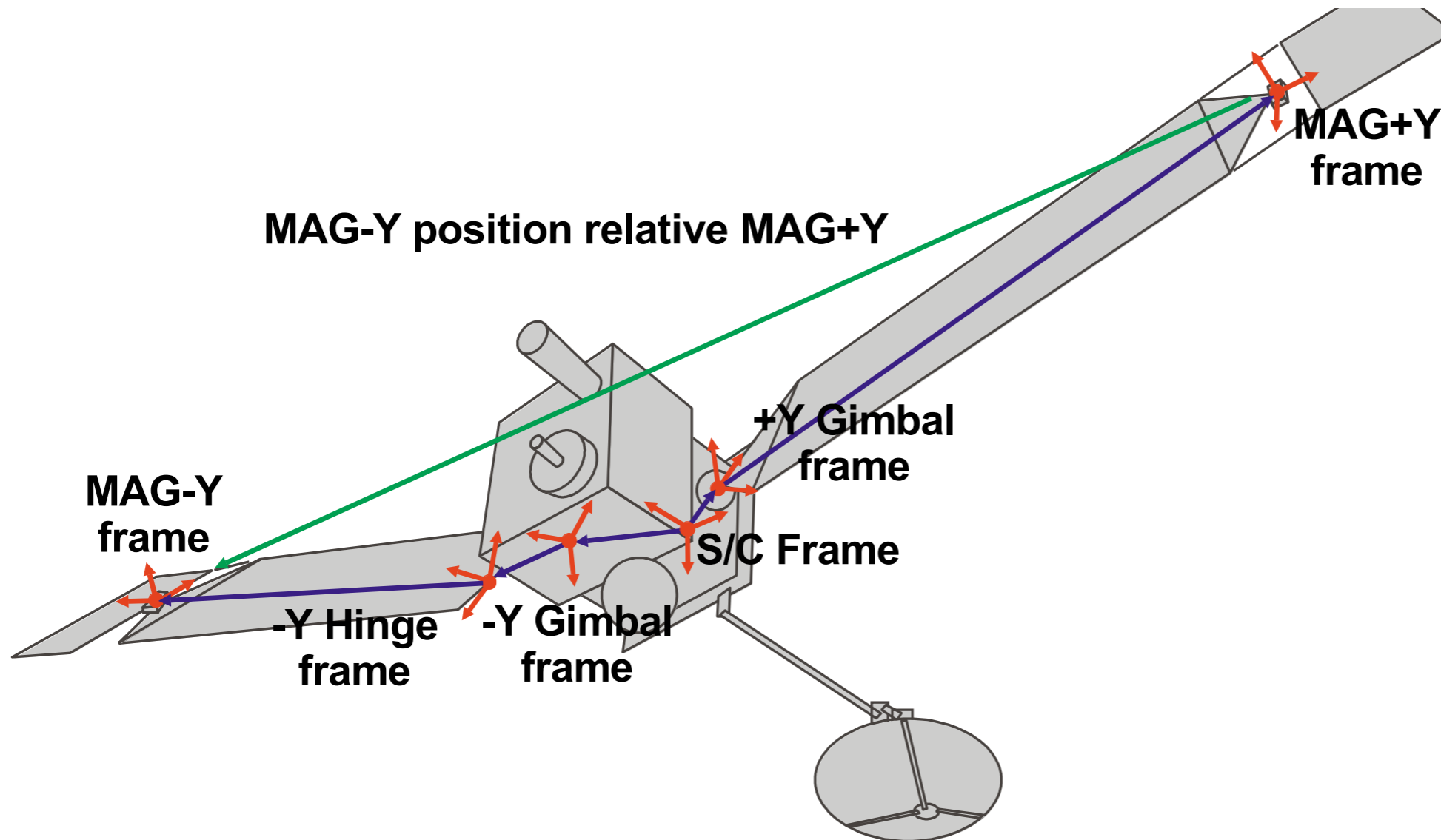
Example 2



Compute solar azimuth and elevation at the Huygens probe landing site

```
CALL SPKPOS ('SUN', ET, 'HUYGENS_LOCAL_LEVEL', 'LT+S', 'HUYGENS_PROBE', SUNVEC, LT)
CALL RECLAT(SUNVEC, R, AZIMUTH, ELEVATION)
ELEVATION = -ELEVATION
IF (AZIMUTH .LT. 0.D0) THEN
    AZIMUTH = AZIMUTH + TWOPI()
ENDIF
```

Example 3

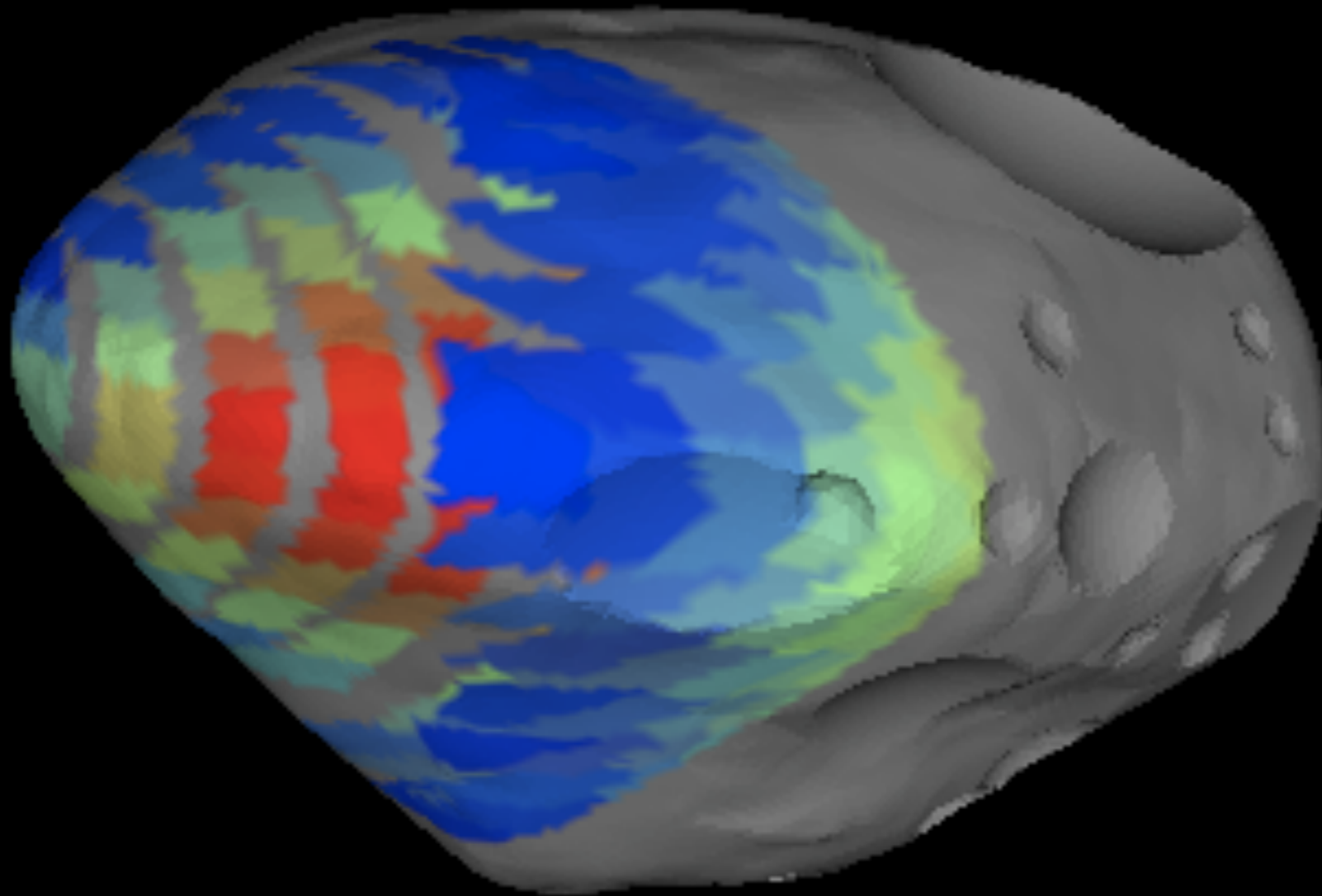


Find the position of one MGS MAG sensor with respect to the other in the MGS s/c frame. Also find the relative orientation of the sensors:

```
CALL SPKEZR('MGS_MAG-Y', ET, 'MGS_SPACECRAFT', 'NONE', 'MGS_MAG+Y', STATE, LT)  
CALL PXFORM('MGS_MAG_+Y_SENSOR', 'MGS_MAG_-Y_SENSOR', ET, MAT)
```

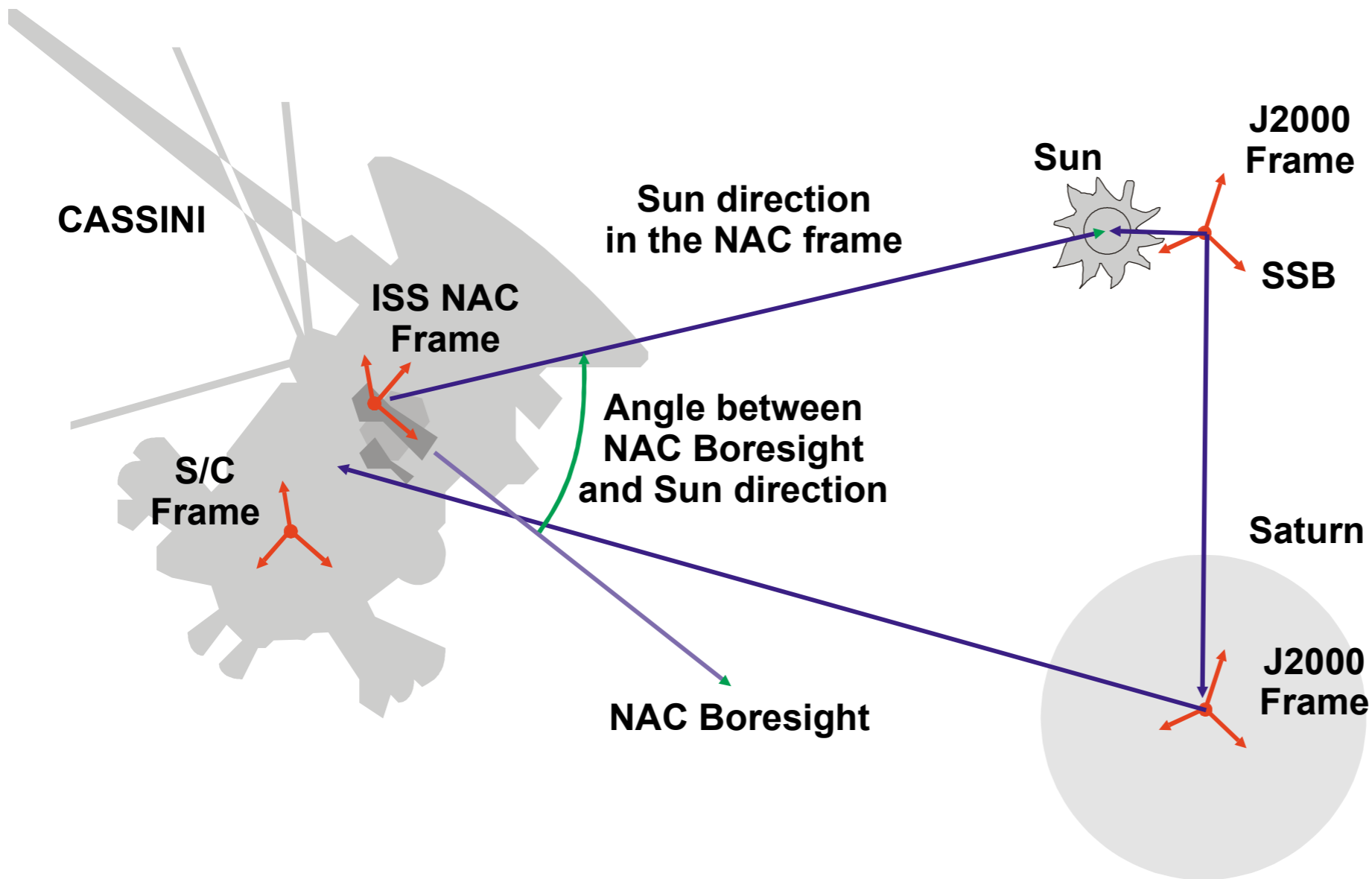
Exemple 4 (officieux)

Steins vu par VIRTIS/OSIRIS sur Rosetta



(Version Alpha)

Inputs/Outputs



Check whether the angle between the camera boresight and the direction to the Sun is within the allowed range:

```
CALL SPKPOS ( 'SUN', ET, 'CASSINI_ISS_NAC', 'LT+S', 'CASSINI', SUNVEC, LT )
```

↑
Target (objet)

↑
Observer (objet)

Gestion des names/IDs

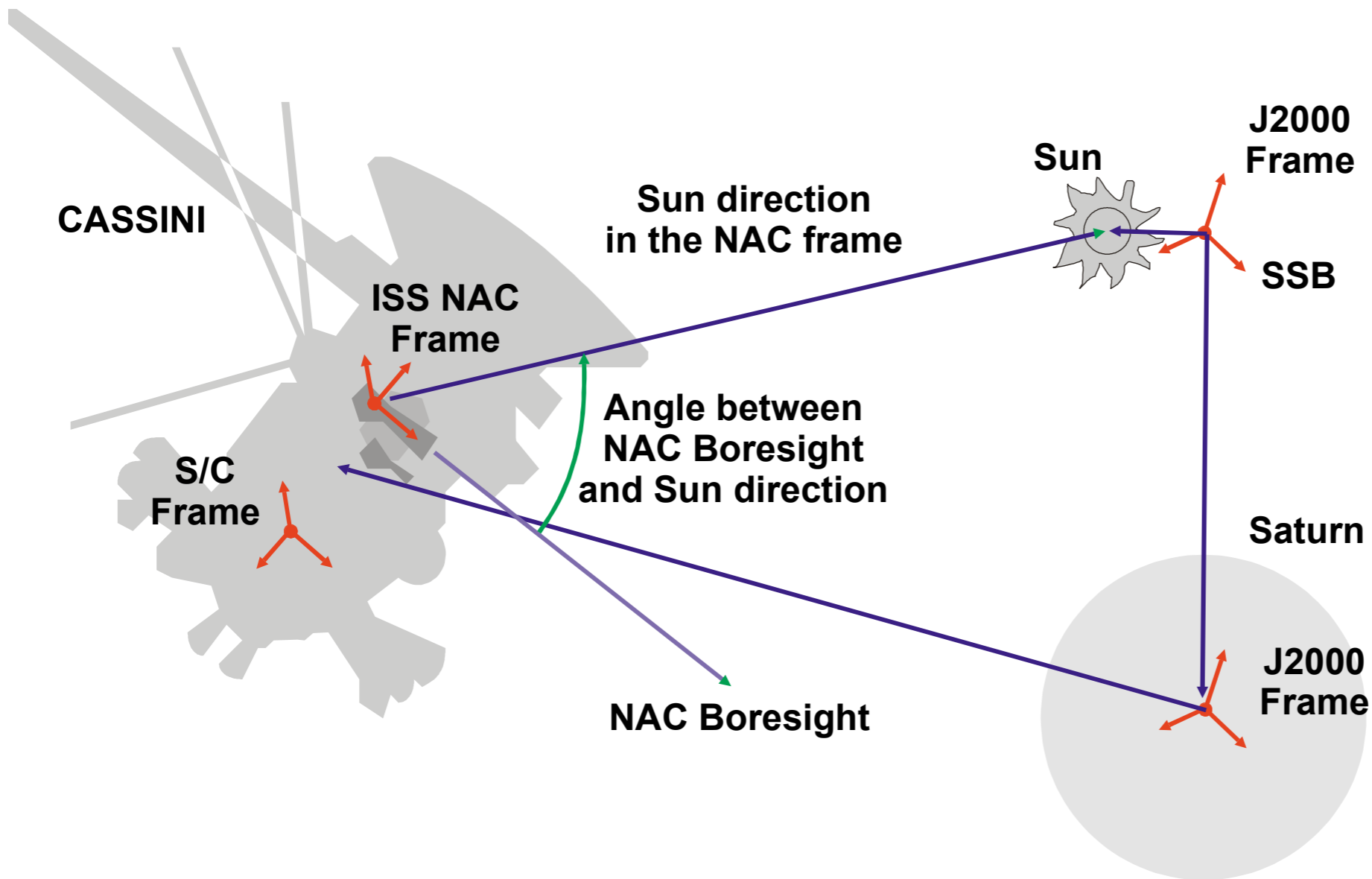
- **SPICE uses IDs and names to identify:**
 - physical objects
 - reference frames
- **An ID is an integer number**
- **A name is a text string**

- **IDs are used primarily as data identifiers inside SPICE kernels**
- **Names are used primarily as input and output arguments in SPICE software interfaces (APIs)**

Examples d'IDs/Names

- **Spacecraft (negative numbers)**
 - Within NASA, this number is generally the negative of the numeric ID assigned by the NASA control authority at GSFC
 - -6 'PIONEER-6', 'P6'
 - -7 'PIONEER-7', 'P7',
 - -82 'CASSINI', 'CAS'
 - -94 'MARS GLOBAL SURVEYOR', 'MGS'
 - ...
- **Sun and Solar System Barycenter (10 and 0)**
 - 0 'SOLAR SYSTEM BARYCENTER', 'SSB'
 - 10 'SUN'
- **Planetary system barycenters (numbers from 1 to 9)**
 - 1 'MERCURY BARYCENTER'
 - 2 'VENUS BARYCENTER'
 - 3 'EARTH MOON BARYCENTER', 'EMB', ...
 - 4 'MARS BARYCENTER'
 - ...
 - 9 'PLUTO BARYCENTER'

Inputs/Outputs



Check whether the angle between the camera boresight and the direction to the Sun is within the allowed range:

```
CALL SPKPOS ( 'SUN' , ET , 'CASSINI_ISS_NAC' , 'LT+S' , 'CASSINI' , SUNVEC , LT )
```

↑
Reference (frame)

Systemes de référence (I)

- **A reference frame is an ordered set of three mutually orthogonal (possibly time dependent) unit-length direction vectors, coupled with a location called the frame's "center" or "origin."**
 - **SPICE documentation frequently uses the shorthand "frame" instead of "reference frame."**
 - **The ordered set of axes of a reference frame is also called a "basis."**
- **A coordinate system specifies the method of locating a point within a reference frame.**
- **A reference frame's center is an ephemeris object whose location is coincident with the origin (0, 0, 0) of the frame.**
 - **The center of the IAU_<body> frame is <body>.**
 - **The center of any inertial frame is (in SPICE) the solar system barycenter.**

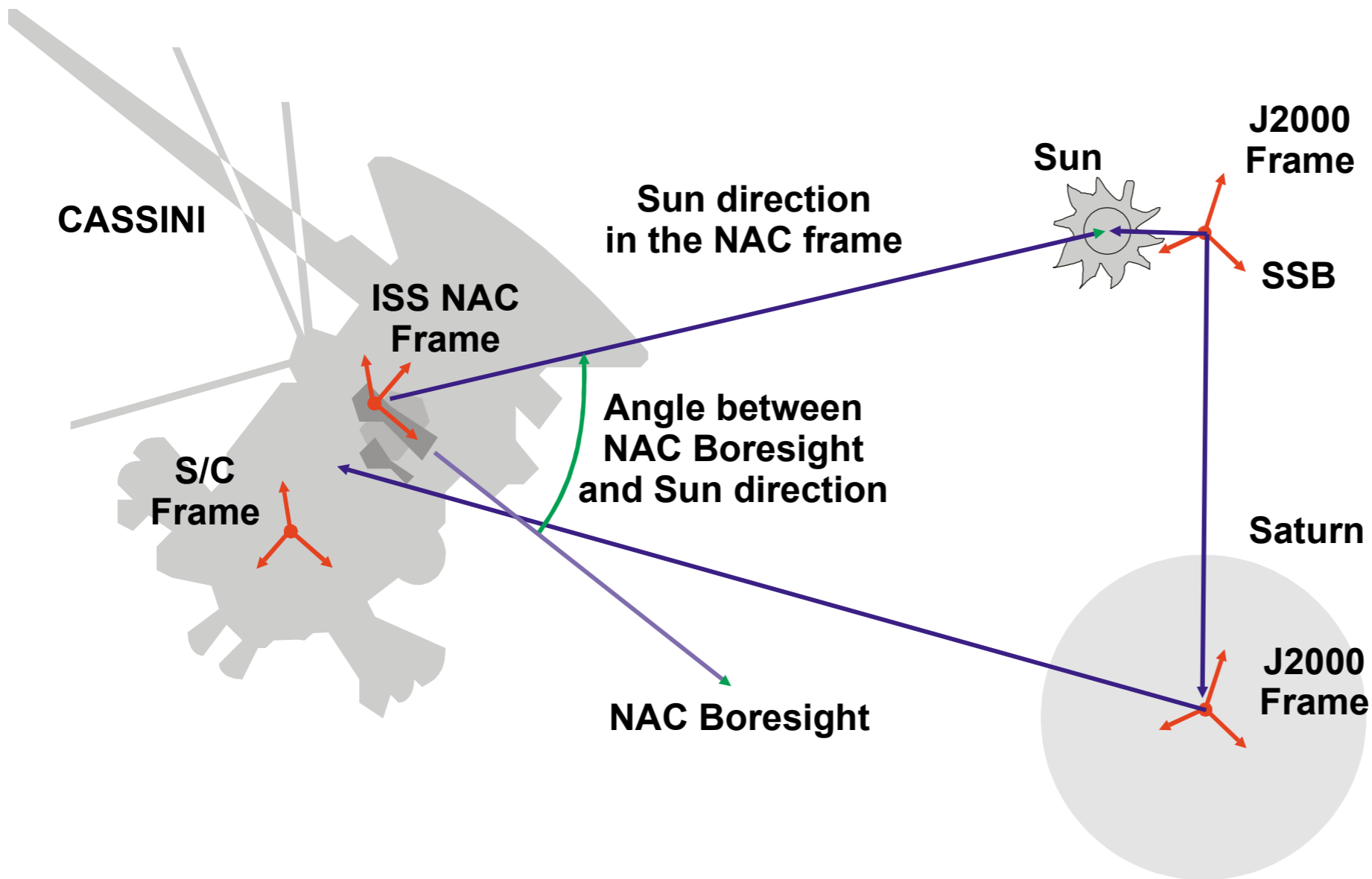
Systemes de référence (II)

- **Inertial**
 - **Non-rotating**
 - » **With respect to fixed stars**
 - **Non-accelerating origin**
 - » **Velocity is typically non-zero; acceleration is negligible**
 - **Examples:**
 - » **J2000 (also called ICRF), B1950**
- **Non-Inertial**
 - **Examples**
 - » **Body-fixed**
 - **Centered at body center**
 - **Topocentric**
 - » **Instrument**
 - » **Dynamic frames**
 - **For example, frames defined by time-dependent vectors**

Systemes de référence (III)

- **Inertial frames (positive integers starting at 1)**
 - 1 'J2000'
 - ...
 - 16 'MARSIAU'
 - 17 'ECLIPJ2000'
 - ...
- **Body-fixed frames (positive integers starting at 10001)**
 - 10001 'IAU_MERCURY_BARYCENTER'
 - ...
 - 10011 'IAU_MERCURY'
 - ...
 - 10020 'IAU_MOON'
 - ...
 - 10081 'EARTH_FIXED'
 - ...

Inputs/Outputs



Check whether the angle between the camera boresight and the direction to the Sun is within the allowed range:

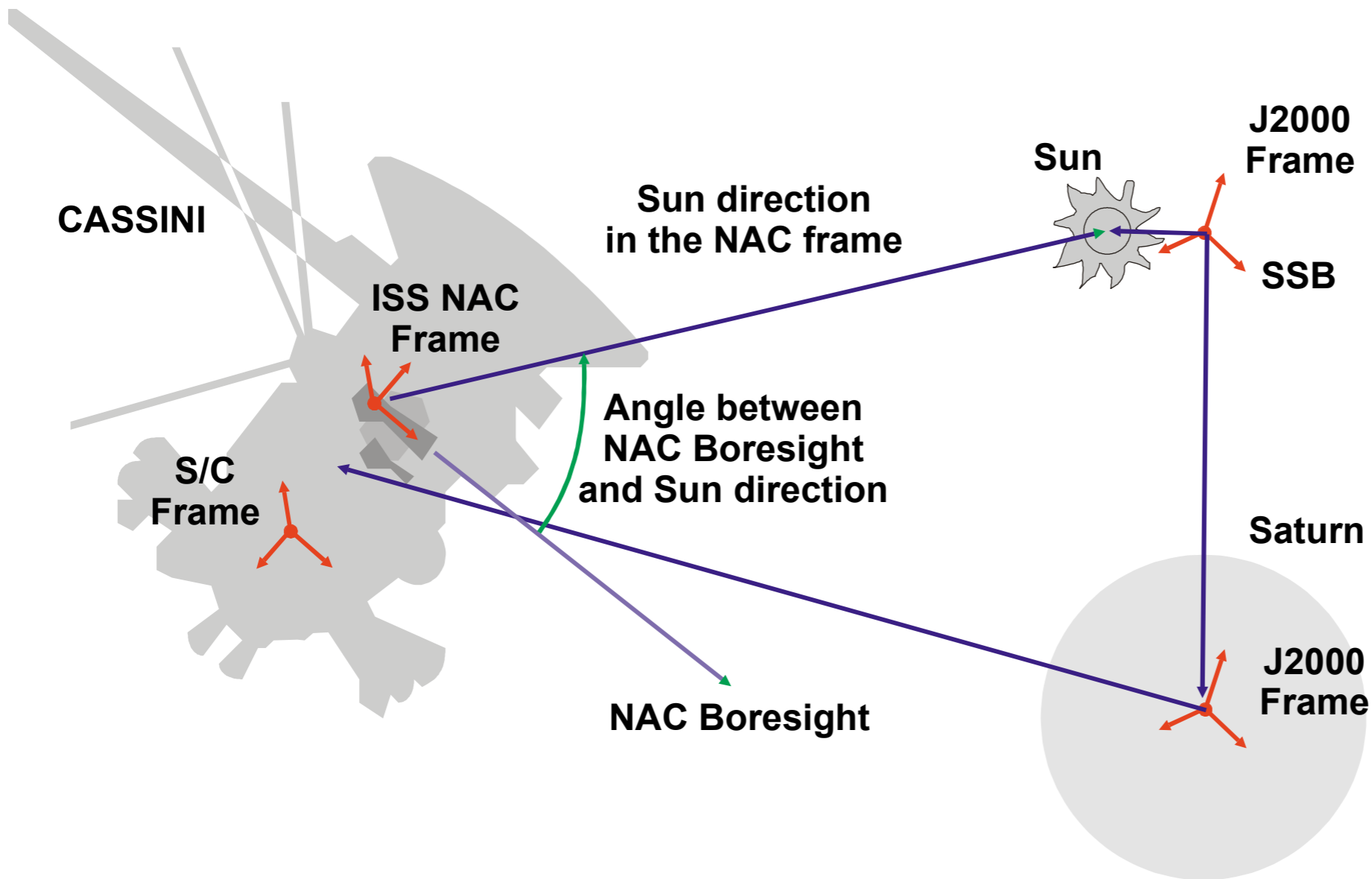
```
CALL SPKPOS ( 'SUN' , ET , 'CASSINI_ISS_NAC' , 'LT+S' , 'CASSINI' , SUNVEC , LT )
```

↑
(Ephemeris) Time

Gestion du temps

- Time inputs and outputs in users' SPICE-based programs are usually **strings** representing epochs in these three time systems:
 - Coordinated Universal Time (**UTC**)
 - Spacecraft Clock (**SCLK**)
 - Ephemeris Time (**ET**, also referred to as Barycentric Dynamical Time, **TDB**)
- Time stamps in kernel files, and time inputs and outputs to SPICE routines reading kernel data and computing derived geometry, are double precision **numbers** representing epochs in these two time systems:
 - Numeric Ephemeris Time (TDB), expressed as ephemeris seconds past J2000
 - Encoded Spacecraft Clock, expressed as clock ticks since the clock start
- SPICE provides routines to perform conversions between string and numeric times using data from these two kernels:
 - Leapseconds Kernel (LSK) containing data for UTC \Leftrightarrow ET conversion
 - Spacecraft Clock Kernel (SCLK) containing data for ET \Leftrightarrow SCLK conversion

Inputs/Outputs

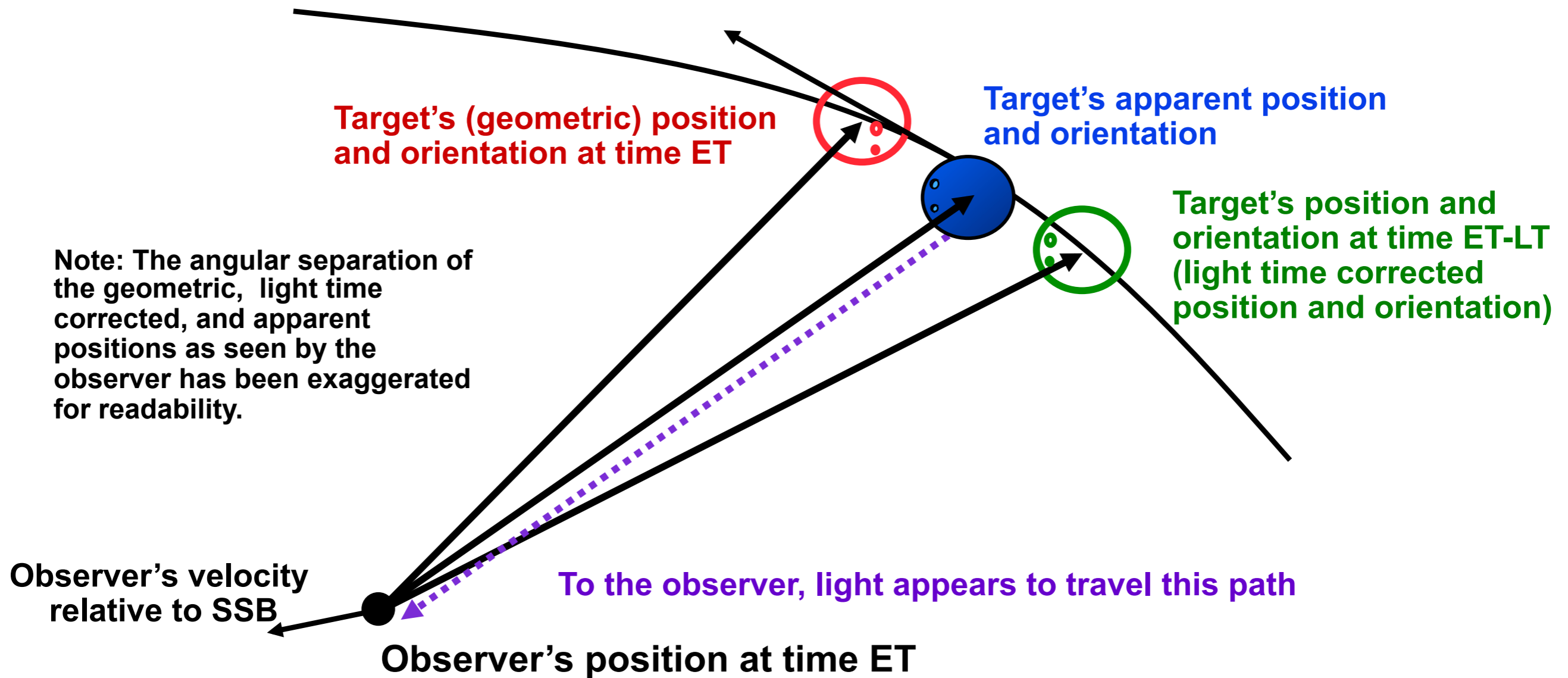


Check whether the angle between the camera boresight and the direction to the Sun is within the allowed range:

```
CALL SPKPOS ( 'SUN' , ET , 'CASSINI_ISS_NAC' , 'LT+S' , 'CASSINI' , SUNVEC , LT )
```

↑
Light Time + Stellar aberration-corrections

Correction position apparent/réelle



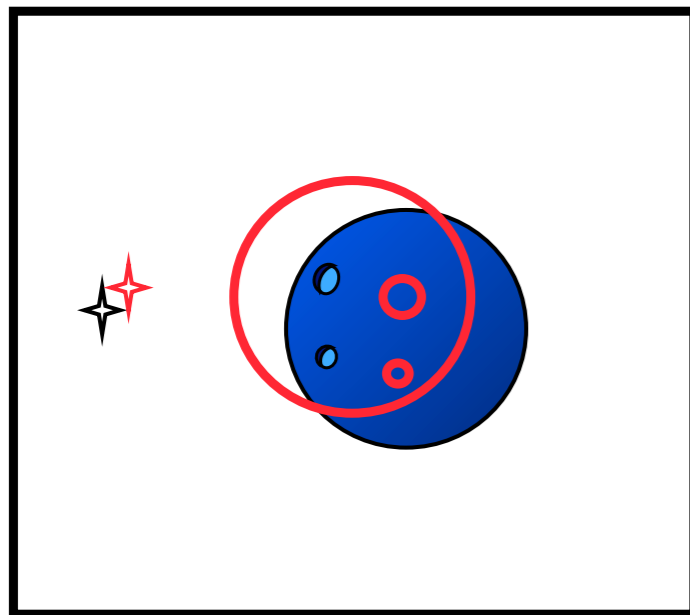
At time ET, the observer's camera records photons emitted from the target at time ET-LT, where LT is the one-way light time.

The vector from the observer at ET to the location of the target at ET-LT is displaced by a physical phenomenon called stellar aberration. The displaced vector yields the apparent position of the target.

Correction position apparent/réelle

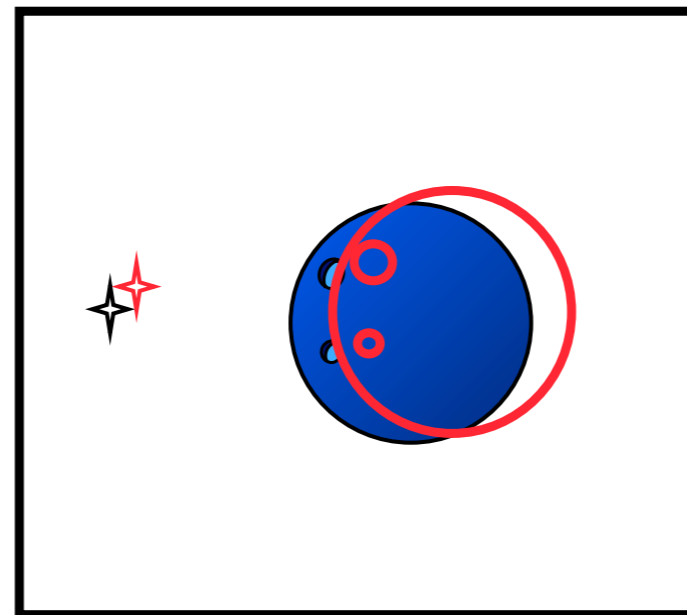
We compare the predicted appearance of a photograph from an optical camera against the actual photograph. We show three predictions derived using different aberration corrections: NONE, LT ("light time only"), and LT+S ("light time plus stellar aberration").

For each prediction, we use red overlays to indicate the expected location in the photo of the images of an extended target body (for example, a natural satellite), of features on the surface of the target body, and of a star.



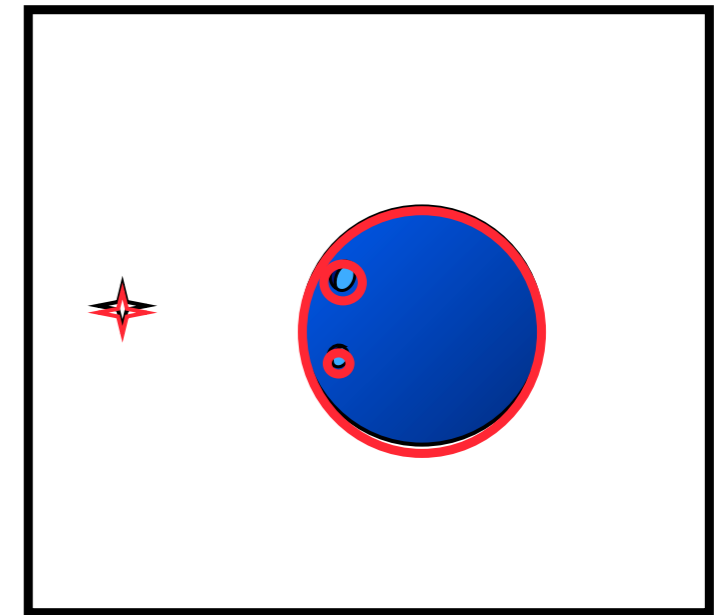
NONE

Predicted images using uncorrected target position and orientation and uncorrected star direction vector



LT

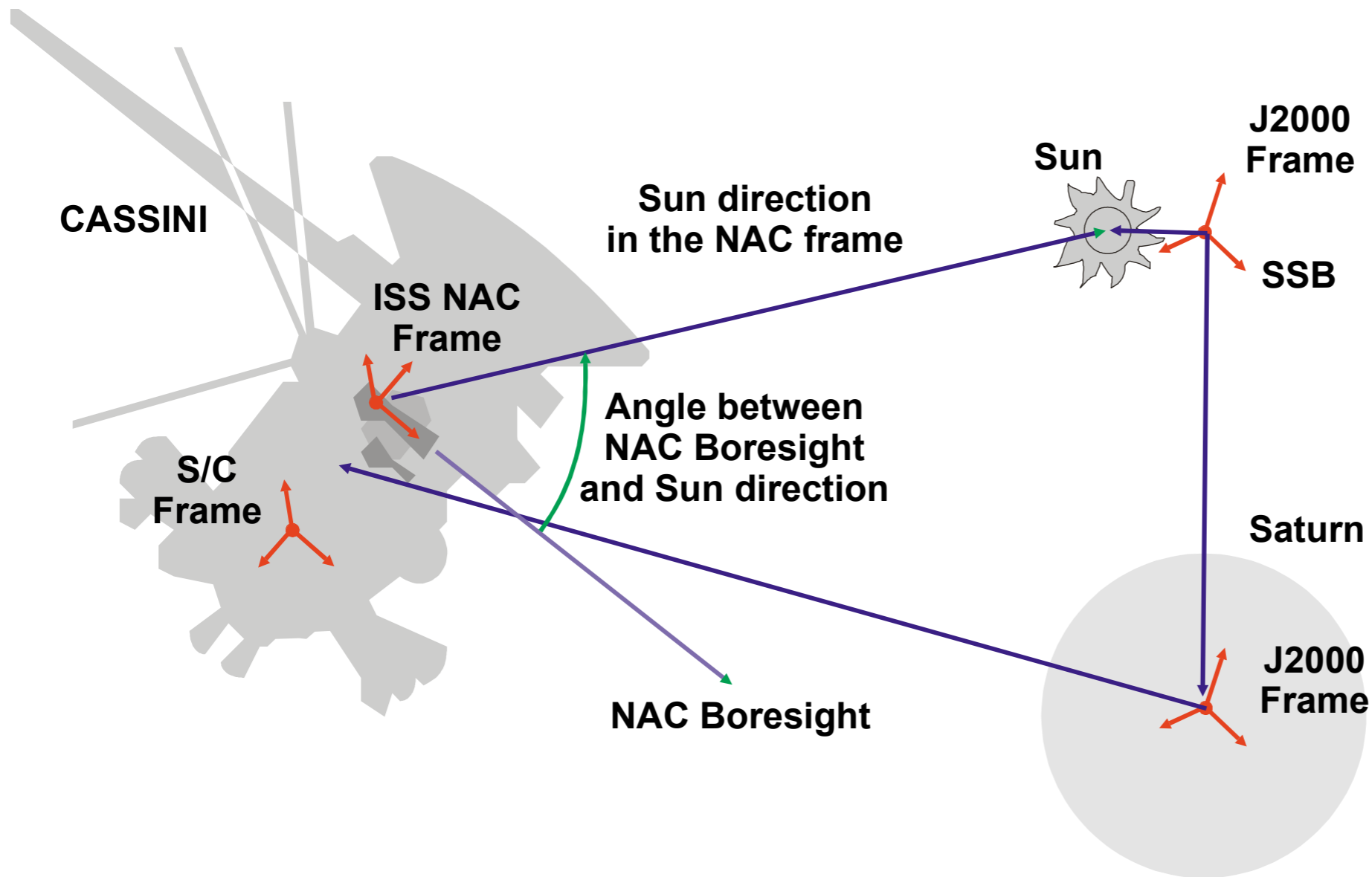
Predicted images using light time-corrected target position and orientation and uncorrected star direction vector



LT+S

Predicted images using light time and stellar aberration-corrected target position, light time-corrected target orientation, and stellar aberration-corrected star direction vector

Fichiers de données (kernels)



Required Kernels:

- Generic LSK
- Mission FK
- Spacecraft SCLK
- Camera IK
- Planetary Ephemeris SPK
- Spacecraft SPK
- Spacecraft CK

Check whether the angle between the camera boresight and the direction to the Sun is within the allowed range:

```
CALL SPKPOS ( 'SUN' , ET , 'CASSINI_ISS_NAC' , 'LT+S' , 'CASSINI' , SUNVEC , LT )
```

Use the FURNISH routine to load all kernels—text and binary.

```
CALL FURNISH ( 'name.ext' )
```

Types de kernels (I)



- **Space vehicle ephemeris (trajectory)**
- **Planet, satellite, comet and asteroid ephemerides**
- **More generally, position of something relative to something else**

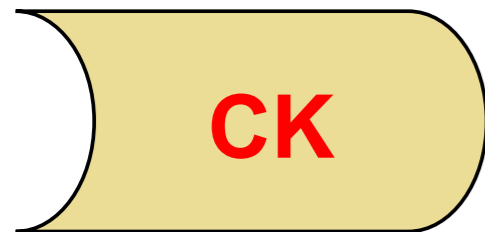


- **Planet, satellite, comet and asteroid orientations, sizes, shapes**
- **Possibly other similar “constants” such as parameters for gravitational model, atmospheric model or rings model**



- **Instrument field-of-view size, shape, orientation**
- **Possibly additional information, such as internal timing**

Types de kernels (II)



- **Instrument platform (e.g. spacecraft) attitude**
- **More generally, orientation of something relative to a specified reference frame**



- **“Events,” broken into three components:**
 - **ESP: Science observation plans**
 - **ESQ: Spacecraft & instrument commands**
 - **ENB: Experiment “notebooks” and ground data system logs**

EK is not much used

Types de kernels (III)

FK

- **Frames**
 - Definitions of and specification of relationships between reference frames (coordinate systems)
 - Both “fixed” and “dynamic” frames are available

LSK

- **Leapseconds Tabulation**
 - Used for UTC <--> TDB (ET) time conversions

SCLK

- **Spacecraft Clock Coefficients**
 - Used for SCLK <--> TDB (ET) time conversions

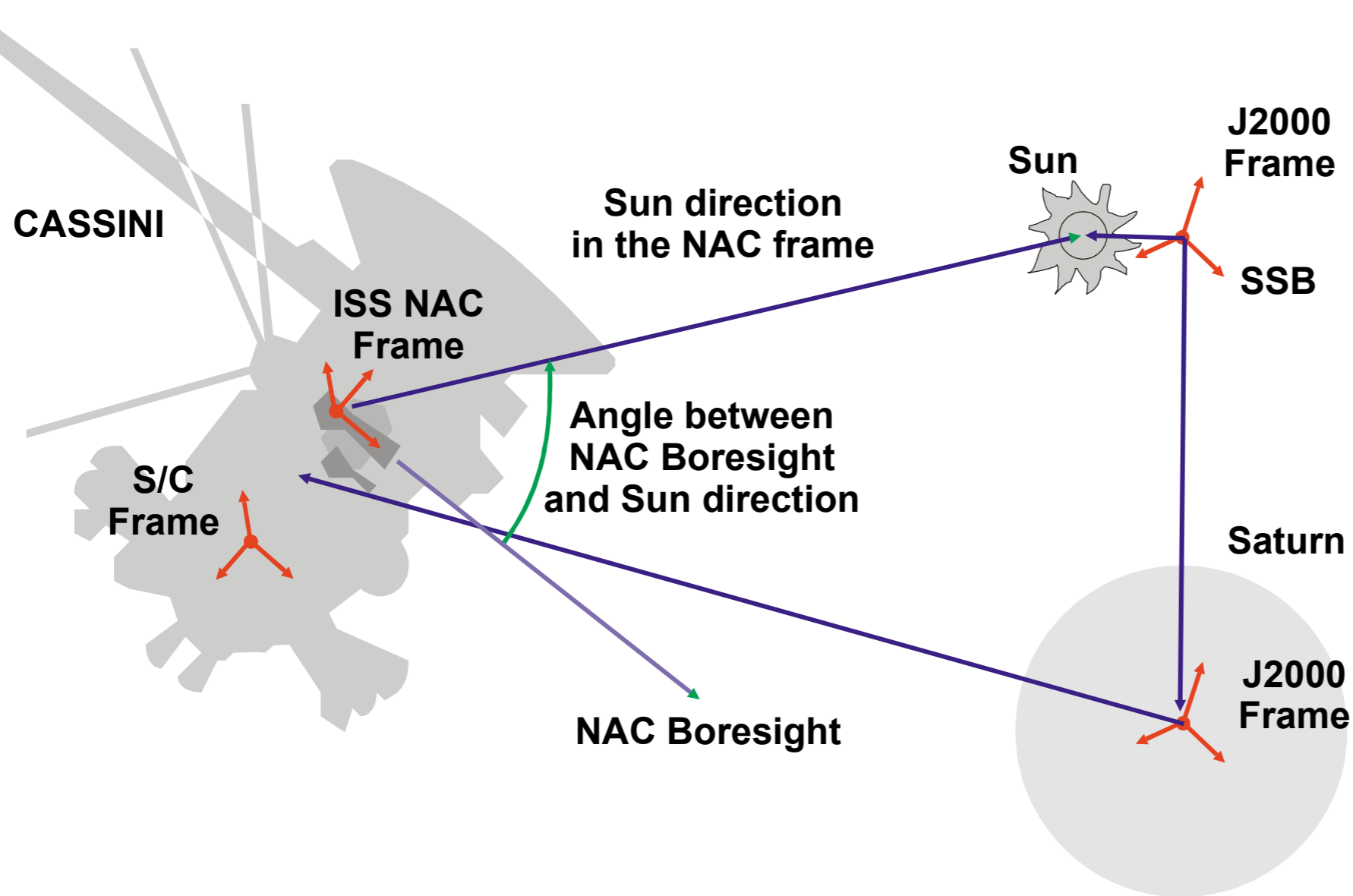
**Under
Development**

- **Shape models (DEM and tessellated plates) (DSK) ¹**
- **Star (sky) catalog ²**

¹ under development now

² development is stalled

Example 1



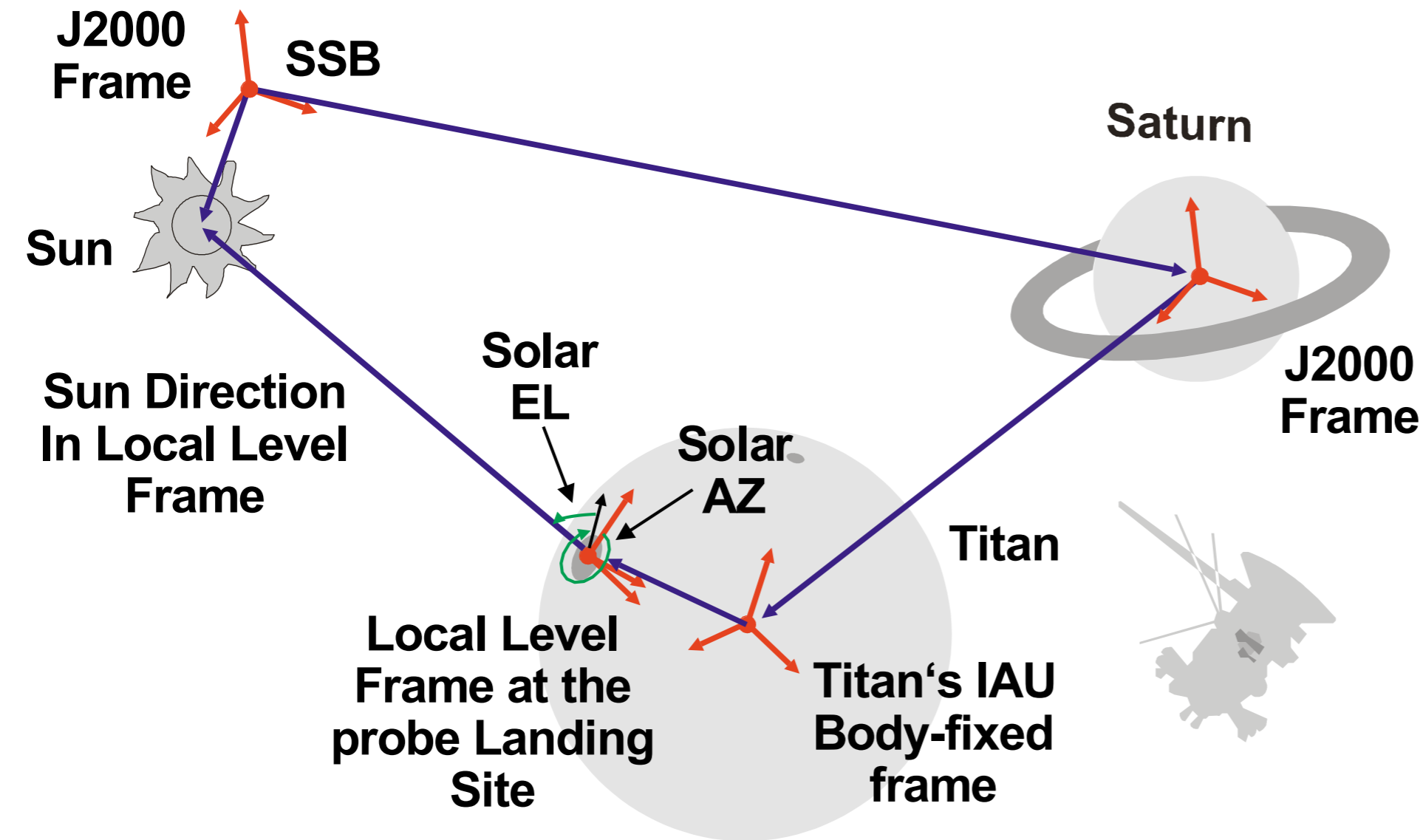
Required Kernels:

- Generic LSK
- Mission FK
- Spacecraft SCLK
- Camera IK
- Planetary Ephemeris SPK
- Spacecraft SPK
- Spacecraft CK

Check whether the angle between the camera boresight and the direction to the Sun is within the allowed range:

```
CALL SPKPOS ( 'SUN', ET, 'CASSINI_ISS_NAC', 'LT+S', 'CASSINI', SUNVEC, LT )
ANGLE = VSEP ( NAC_BORESIGHT_nac, SUNVEC )
IF ( ANGLE .LE. CONSTRAINT ) WRITE (*,*) 'WE ARE IN TROUBLE!'
```

Exemple 2



Required Kernels:

- Generic LSK
- Generic PCK
- Mission FK
- Planetary Ephemeris SPK
- Satellite Ephemeris SPK
- Landing Site SPK

Compute solar azimuth and elevation at the Huygens probe landing site

```
CALL SPKPOS ('SUN', ET, 'HUYGENS_LOCAL_LEVEL', 'LT+S', 'HUYGENS_PROBE', SUNVEC, LT)
CALL RECLAT (SUNVEC, R, AZIMUTH, ELEVATION)
ELEVATION = -ELEVATION
IF (AZIMUTH .LT. 0.D0) THEN
    AZIMUTH = AZIMUTH + TWOPI ()
ENDIF
```

En pratique

- **Software** <http://naif.jpl.nasa.gov/naif/toolkit.html>
 - Subroutine libraries, with source code
 - » **SPICELIB (Fortran)**
 - » **CSPICE (C)**
 - » **Icy (C)**
 - » **Mice (C and Matlab script)**
 - Executable programs
 - » **application and utility programs**
 - » **cookbook examples**
 - Installation/build scripts
- **Documentation** <http://naif.jpl.nasa.gov/naif/tutorials.html>
 - Available in ASCII and HTML
- **Data** <http://naif.jpl.nasa.gov/naif/data.html>
 - kernel files

Question?

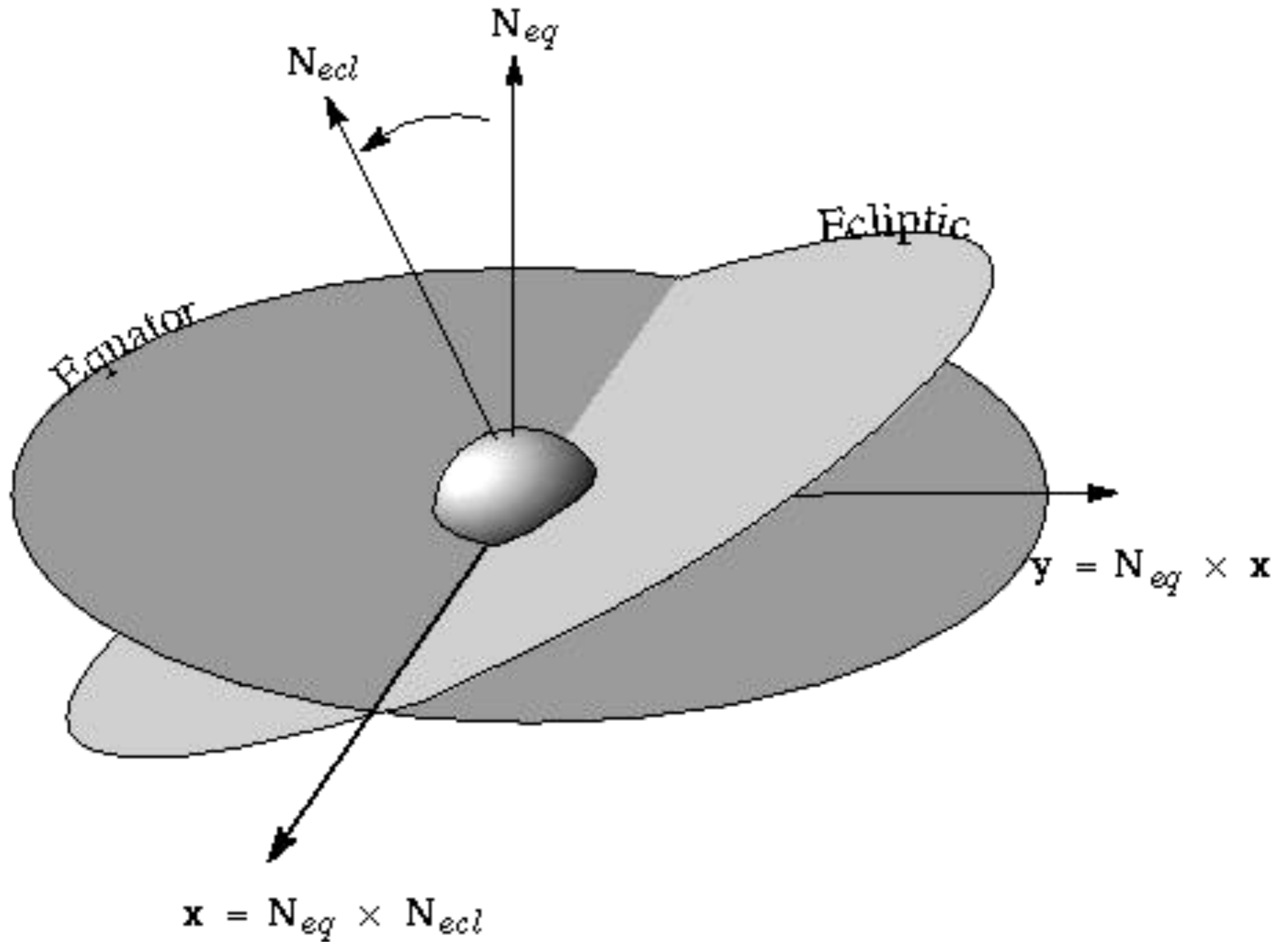
Précision?

Point obscure?

Suggestion?...



J2000 (ICRF)



Format et extension des kernels

SPK:

.bsp binary SPK file

PCK:

.tpc text PCK file
(The most common type Pck)

.bpc binary PCK file
(Very few instances of this)

IK:

.ti text IK file

FK:

.tf text FK file

LSK:

.tls text LSK file

CK:

.bc binary CK file

SCLK:

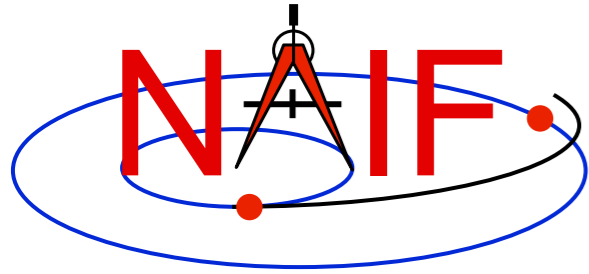
.tsc text SCLK file

MK:

.tm text meta-kernel file (“FURNISH kernel”)

DSK:

.bds binary DSK file



Barycentric Dynamical Time

Navigation and Ancillary Information Facility

- **Barycentric Dynamical Time (TDB) and Ephemeris Time (ET) are synonyms in SPICE documentation.**
- **TDB is**
 - a mathematical Ideal used in the equations of motion.
 - used as the independent time variable for many SPICE subroutine interfaces.
 - related to ideal time at the Solar System Barycenter (TCB) by a scale factor, so that TDB advances on average at the same rate as TAI.

Les fichiers de données (kernels)

- **SPK**
 - Spacecraft and Planet Ephemeris
- **PcK**
 - Planetary Constants, for natural bodies
 - » Orientation
 - » Size and shape
- **IK**
 - Instrument
- **CK**
 - Pointing (“C-matrix”)
- **EK**
 - Events, up to three distinct components
 - » ESP: science plan
 - » ESQ: sequence
 - » ENB: experimenter’s notebook
- **FK**
 - Reference frame specifications
- **SCLK**
 - Spacecraft clock correlation data
- **LSK**
 - Leapseconds
- **Meta-Kernel** (a.k.a. “FURNSH kernel”)
 - Mechanism for aggregating and easily loading a collection of kernel files
- **DSK (under development)**
 - Digital shape kernel
 - » Tessellated plate model
 - » Digital elevation model

- **SCLK string formats vary from one spacecraft clock to the next.**
 - **Cassini: Maximum reading for partition 1 = 1/4294967295.255**
 - » **Partition number: 1**
 - » **Seconds: 4294967295**
 - » **Ticks (for Cassini, unit = 1/256 second): 255**
 - **Galileo: Maximum reading for partition 1 = 1/16777215:90:09:07**
 - » **Partition number: 1**
 - » **"RIM" count (unit = 60 2/3 seconds): 16777215**
 - » **"Mod 91" count (unit = 2/3 second): 90**
 - » **"RTI" count (unit = 1/15 second): 9**
 - » **"Mod 8" count (unit = 1/120 second): 7**
- **The format of spacecraft clock and the relationship between tick count and other time systems (usually UTC) is captured in a SPICE SCLK kernel**